

The singularity theorems of General Relativity and their low regularity extensions

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based on past and ongoing work mainly with

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The singularity theorems in General Relativity

About

- rigorous results in Lorentzian differential geometry
- physically reasonable ass. (collapse, expanding univ.) \leadsto singularities
- singularity \sim incomplete causal geodesic
- occurrence of spacetime singularities as a generic feature of GR

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Here: Explain the structure of arguments & **low regularity** extensions

- **causality theory** for metrics down to g (loc. Lipschitz) continuous
- **analysis of geodesic focusing** down to $g \in C^1$, or locally Lipschitz

The structure of the singularity theorems

Pattern theorem (Senovilla, 1998)

A spacetime (M, g) is singular if it satisfies:

- (I) A suitable initial condition,
- (E) a condition on the curvature (energy condition),
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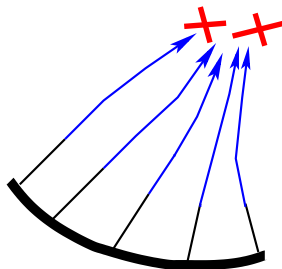
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- (I) \leadsto causal geodesics start focusing
- (E) \leadsto focusing goes on (Raychaud. Riccati)
 - \leadsto focal/conjugate point
 - \leadsto geos. stop maximising
- (C) \leadsto there are maximising causal geos

Resolution: some causal geodesics
stop existing before (first) conjugate point



The three classical theorems

Thm. (Penrose, 1965) A spacetime is future null incomplete if

(E) $\text{Ric}(X, X) \geq 0$ for all X null (NEC)

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Thm. (Hawking and Penrose, 1970) Improvements:

- (E) (SEC) and genericity
- (C) chronological (**only!**)
- (I) cp. achronal set w/o. edge, trapped surface/point/submanifold

Analytic core: geodesic focusing under curvature bds.

- Geodesic curves $\gamma : I \rightarrow M$ with $\nabla_{\dot{\gamma}}\dot{\gamma} = 0$
- For data $p \in M$, $v \in T_pM$ unique maximally extended solution on $I_{p,v}$

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Message: Bounds on Ricci curvature bound length of maximisers.

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- causal geodesic completeness \implies inextendability
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 - apply thms. to maximally extended spacetimes

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!! Here **regularity of extensions** becomes an issue
classical thms. only rule out C^2 -extensions
but extensions are ok as long as field equations make sense

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- alternative settings
 - causal extensions [Minguzzi,19]
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Main challenges

- (1) extend causality theory: deprived of main geometric tools
- (2) extend analytic core: focusing under distributional curvature bds.

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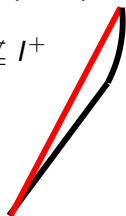
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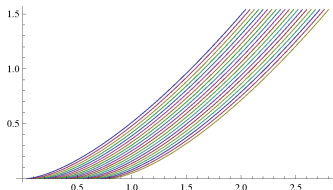
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(1) push up principle fails

$$(J^+ \circ I^+) \not\subseteq I^+$$



(2) light cones bubble up



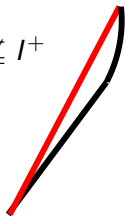
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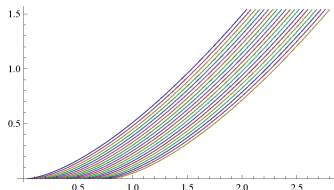
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- ✓ some topological features are more robust: Avez-Seifert [Sämman, 16]
- cone structures [Bernard & Suhr, 18], [Minguzzi, 19]

Key analytic techniques: convolution & Friedrichs lemma

Basis: chartwise regularisation of metric by convolution

$$g_\varepsilon(x) := g \star_M \rho_\varepsilon(x) := \sum \chi_i(x) \psi_i^* \left((\psi_{i*}(\zeta_i \cdot g)) * \rho_\varepsilon \right)(x).$$

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Lemma (Regularisation for Lipschitz g) [CG,12], [CGHKS,24]

There are smooth Lorentzian metrics $\check{g}_\varepsilon, \hat{g}_\varepsilon$ with

- $\check{g}_\varepsilon \prec g \prec \hat{g}_\varepsilon$ sandwiched light cones via tweaked convolution
- $\check{g}_\varepsilon, \hat{g}_\varepsilon \rightarrow g$ in $W_{\text{loc}}^{1,p}(M)$ ($1 \leq p < \infty$)

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Non-smooth focusing: The rough guide

- 1 Formulate distributional (EC) for $g \in \text{Lip}$
- 2 Derive surrogate (EC) for $\text{Ric}[\check{g}_\varepsilon](X, X)$ (on K cp.) ...
- 3 still show smooth focusing for \check{g}_ε -geodesics ...extend analysis
- 4 show that geos of g stop maximising/existing ...extend causality

Non-smooth focusing: Following the rough guide

- Start with distributional condition $\text{Ric}[g] \geq 0$
- Problem: $\text{Ric}[\check{g}_\varepsilon] \rightarrow \text{Ric}[g]$ only distributionally
 \leadsto cannot carry $\text{Ric}[g] \geq 0$ through construction

- Solution: compatibility of distinct regularisations

$$\text{Ric}[\check{g}_\varepsilon] - \underbrace{\text{Ric}[g] \star_M \rho_\varepsilon}_{\geq 0} \quad \text{locally uniformly only for } g \in C^1$$

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Lemma (Compatibility of regularisations for $g \in \text{Lip}$) [CGHKS,24]

- $\|\text{Ric}[\check{g}_\varepsilon] - \text{Ric}[g] \star_M \rho_\varepsilon\|_{L^p(K)} \rightarrow 0$ ($1 \leq p < \infty$)
- $\|\text{Ric}[\check{g}_\varepsilon] - \text{Ric}[g] \star_M \rho_\varepsilon\|_{L^\infty(K)} \leq C_K$ (advanced Friedrichs lemma)

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Feed into the geometric focusing machinery via

- volume estimates [Treude & Grant, 13]
- segment inequality [Graf, Kontou, Ohanyan, Schinnerl, 24]
estimate line- \int by volume- \int inspired by [Cheeger & Colding, 96]

The Penrose and the Hawking theorems in C^1 /Lipschitz

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Thm. (CGHKS, 2024) Let g be a **loc. Lipschitz** metric s.t.

- (E) $\text{Ric}(X, X) \geq 0$ for all X timelike **as distribution (DSEC)**
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- (I) **slab mean curvature bounded below by $b > 0$.**

then the future time separation τ_Σ from Σ is bounded by $\tau_\Sigma \leq \frac{1}{b}$.

The Hawking-Penrose theorem in C^1

Thm. (Hawking, Penrose, 1970) A C^2 -spacetime is caus. incompl. if

(E) (SEC) and the genericity cond. hold i.e.,

$$\forall \gamma \quad \exists t_0 : \quad R(., \dot{\gamma})\dot{\gamma} : [\dot{\gamma}(t_0)]^\perp \rightarrow [\dot{\gamma}(t_0)]^\perp \quad \text{non-trivial} \quad (\text{GC})$$

(C) it is chronological

and it has at least one of

(I1) a cp. achronal set without edge, (I2) a future trapped surface

(I3) a closed future trapped submanifold w. $\sum_{i=1}^{n-m} \langle R(E_i, \dot{\gamma})\dot{\gamma}, E_i \rangle \geq 0$

(I3) a future trapped point .

The Hawking-Penrose theorem in C^1

Thm. (KOSS, 2022) A C^1 -spacetime is caus. incompl. if

(E) (DSEC) and the \mathcal{D}' -genericity cond. hold i.e.,

$$\forall \gamma \quad g(R(\tilde{X}, \tilde{V})\tilde{V}, \tilde{X}) \geq C > 0 \text{ in } \mathcal{D}'^{(1)}(U) \quad (\text{DGC})$$

(C) it is causal

(B) it is maximally, causally non-branching and it has at least one of

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(I3) a future trapped point all in the support sense.

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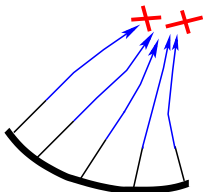
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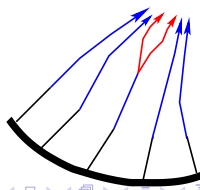
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(I3) a future trapped point all in the support sense.



or



Summary, discussion, results beyond the spacetime setting

(0) spacetime results above for $g \in C^1/\text{Lipschitz}$

- close to the class. results; natural extensions of most notions
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Summary, discussion, results beyond the spacetime setting

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(1) causal cone structures [Minguzzi, 2019]

- upper semi-cont. distribution of cones on M (generalises light cones)
- causal core of singularity theorems may be established
- analytic parts (ECs) only to produce sets with specific causality props.

(Causal Penrose theorem, Minguzzi, 2019)

Let (M, C) be a globally hyperbolic closed cone structure admitting a non-compact stable Cauchy hypersurface Σ . Then there are no compact future trapped sets and if Σ is non-empty and compact there is an inextendible future null geodesic entirely contained in $E^+(S)$.

(2) Synthetic approaches: Lorentzian length spaces

- causal space $(X, d, \leq\ll, \tau)$ with τ intrinsic
- (timelike) sectional curvature bounds via triangle comparison

Results beyond the spacetime setting

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Let $Y = (a, b) \times_f X$ be a warped product (X metric length space, $f \in C^\infty$, non-const.) with positive timelike sectional curvature. Then $a > -\infty$ or $b < \infty$ and hence Y is past/future timelike geodesically incomplete.

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$$L(\gamma = (\alpha, \beta)) = \int \sqrt{\dot{\alpha}^2 - (f \circ \alpha)^2 v_\beta^2} \quad \text{with } v_\beta \text{ the metric derivative.}$$

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- smooth metric measure spacetimes [McCann, 20], [Mondino, Suhr, 22]
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Let X be a timelike non-branching, globally hyperbolic LLS with a TMCP property. Let V be a Borel achronal future timelike complete subset with mean curvature bded above. Then every future timelike geodesic starting in V has a bounded maximal domain of existence.

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- synthetic (NEC) [McCann, 23], [Braun, McCann, 24]
- first-order Sobolev calculus on metric measure spacetimes
(maximal weak subslope of time functions akin L-modulus of diff.)
[Beran, Braun, Calisti, Gigli, McCann, Ohanyan, Rott, Sämann, 24]

Some Literature

- [CV,24] F. Cavalletti and A. Mondino. Optimal transport in Lorentzian synthetic spaces, synthetic timelike Ricci curvature lower bounds and applications. *Camb. J. Math.*, 12(2):417–534, 2024.
- [GGKS,18] M. Graf, J. Grant, M. Kunzinger, R. Steinbauer, The Hawking-Penrose Singularity Theorem for $C^{1,1}$ -Lorentzian Metrics. *Commun. Math. Phys.*, 360(3): 1009–1042, 2018.
- [G,20] M. Graf, Singularity theorems for C^1 -Lorentzian metrics. *Commun. Math. Phys.*, 378(2):1417–1450, 2020.
- [GKOS,22] M. Graf, E.-A. Kontou, A. Ohanyan, and B. Schinnerl, Hawking-type singularity theorems for worldvolume energy inequalities. *Lett. Math. Phys.* 2024.
- [KOSS,22] M. Kunzinger, A. Ohanyan, B. Schinnerl, R. Steinbauer, The Hawking- Penrose Singularity Theorem for C^1 -Lorentzian Metrics. *Commun. Math. Phys.*, 391:1143–1179, 2022.
- [S,23] R. Steinbauer, The singularity theorems of General Relativity and their low regularity extensions. *Jb. Dt. Math.-Ver.* 125(2), 73–119, 2023