

Synthetic curvature for GR and beyond

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excellent = austria

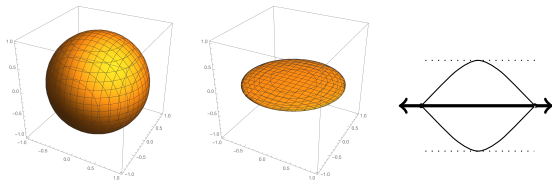
Curvature beyond smooth spacetimes

Why at all?

- physically relevant *models* (matched spacetimes, impulsive wave, etc.)
- *PDE* point-of-view
- *singularities* vs *curvature blow-up* — *CCH* of Penrose
- approaches to *Quantum Gravity* (no metric, e.g. causal sets)

Why it matters?

Basic geometric properties change even if $g \in C^{1,\alpha}$



Squeezing a sphere:

Equator still geodesic
but it's always shorter to
deviate into hemispheres
(Hartman-Wintner '52)

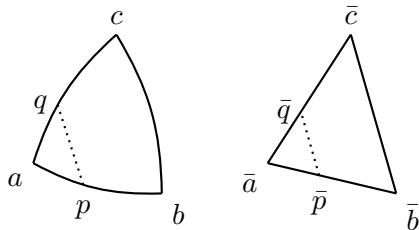
How to detect curvature: A glimpse on Riemannian world

$$\text{Sectional curvature } \text{Sec}(X, Y) = \frac{\langle R(X, Y)Y, X \rangle}{\|X\|^2\|Y\|^2 - \langle X, Y \rangle^2}$$

Theorem (Toponogov) $\text{Sec} \geq K \iff$

For all (small) geodesic triangles $\triangle abc$ in (M, h) consider a comparison triangle $\triangle \bar{a}\bar{b}\bar{c}$ in the 2D space of const. curvature K . Then for all for all p, q on its sides and corresponding comparison points \bar{p}, \bar{q}

$$d_h(p, q) \geq \bar{d}(\bar{p}, \bar{q}).$$



Triangle condition

- needs no manifold structure
- only distances between pts.
- works on metric spaces

Sectional curvature bounds for metric spaces

Definition (Length space)

A metric space (X, d) is called a *length space* if d is intrinsic, i.e.,

$$d(x, y) = \inf\{L(\gamma) \mid \gamma \text{ from } x \text{ to } y \text{ continuous}\}$$

geodesics $\gamma : [0, 1] \rightarrow X$ with $d(\gamma(s), \gamma(t)) = |t - s| \cdot d(\gamma(0), \gamma(1))$

Definition (Synthetic curvature bounds)

A length space has curvature bounded below by K if (locally) for all triangles $\triangle abc$ and their comparison triangles $\triangle \bar{a}\bar{b}\bar{c}$ and all points p, q on its sides and corresponding \bar{p}, \bar{q}

$$d(p, q) \geq \bar{d}(\bar{p}, \bar{q}).$$

curvature bounded below / above: *Alexandrov spaces* / *CAT(K)-spaces*
rich theory since the 1980-ies: GH-convergence, Gromov compactness thm.

How to detect curvature: Lorentzian world

$$\text{Sectional curvature } \text{Sec}(X, Y) = \frac{\langle R(X, Y)Y, X \rangle}{\langle X, X \rangle \langle Y, Y \rangle - \langle X, Y \rangle^2}$$

Kulkarni (1979): If $\text{Sec}(g)$ is bounded below (above), then it is constant.

Definition (“Correct” curvature bounds, Andersson-Howard 1998)

A smooth Lorentzian manifold has $\text{Sec} \geq K$ if *spacelike* sectional curvatures $\geq K$ and *timelike* sectional curvatures $\leq K$.

Theorem (Alexander-Bishop 2008)

A smooth Lorentzian manifold has $\text{Sec} \geq K$ if for all (small) geodesic $\triangle abc$ and their comparison $\triangle \bar{a}\bar{b}\bar{c}$ in 2D space of const. curvature K (Minkowski, (anti-)de Sitter) and all p, q resp. \bar{p}, \bar{q}

$$d_{\text{signed}}(p, q) \geq \bar{d}_{\text{signed}}(\bar{p}, \bar{q}).$$

How to go beyond Lorentzian manifolds?

Riemannian manifolds \subsetneq metric (length) spaces

Lorentzian mfs. / spacetimes \subsetneq ?

What is the analogue of metric (length) spaces in the *Lorentzian setting*?

Serious issue:

- natural analogue to distance: time separation function

$$\tau(p, q) = \sup\{L(\gamma) \mid \gamma \text{ future dir. causal from } p \text{ to } q\}$$

- but triangle inequality is *reversed* \rightsquigarrow no metric structure

\rightsquigarrow *Lorentzian (pre-)length spaces* (Kunzinger-Sämman 2018)

Lorentzian (pre-)length spaces

Causal space: X (metrizable) topological space with *abstract causality*
 \leq preorder on X , \ll transitive relation contained in \leq

Abstract time separation: $\tau: X \times X \rightarrow [0, \infty]$ lower semicontinuous

Definition (Kunzinger-Sämman 2018)

(X, \ll, \leq, τ) is a *Lorentzian pre-length space* if for $p \leq q \leq r$

$$\tau(p, r) \geq \tau(p, q) + \tau(q, r) \quad \text{and} \quad \tau(p, q) \begin{cases} = 0 & \text{if } x \not\ll y \\ > 0 & \Leftrightarrow x \ll y \end{cases}$$

Examples

- *smooth spacetimes* (M, g) with usual time separation function τ
- *Lorentz-Finsler* spacetimes, spacetimes of *low regularity* ($g \in C^0 + \dots$)
- *finite directed graphs* (causal sets)

Lorentzian *causality theory*

τ intrinsic... *Lorentzian length space*

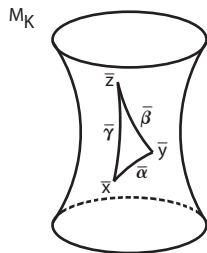
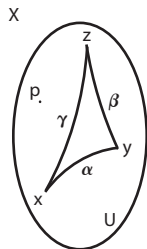
Timelike curvature via triangle comparison

Definition (Synthetic curvature bounds)

(X, \ll, \leq, τ) has *timelike curvature* $\geq K$ if

- 1 some technical conditions hold
- 2 for all *small timelike triangles* Δabc and their comparison $\Delta \bar{a}\bar{b}\bar{c}$ in M_K and all p, q resp. \bar{p}, \bar{q}

$$\tau(p, q) \leq \bar{\tau}(\bar{p}, \bar{q}).$$



Faithful extension of
sectional curvature bounds
to “metric” Lorentzian setting

Selected results

Theorem (Kunzinger-Sämman 2018, Beran-Sämman 2022)

In a strongly causal Lorentzian pre-length space with *timelike curvature bounded below* timelike geodesics do *not branch*.

Theorem (Grant-Kunzinger-Sämman 2019)

A timelike geodesically complete spacetime (or LLS) is *inextendible as a regular LLS*, i.e., any LLS-extension necessarily has unbounded curvature.

Extends (Beem-Ehrlich) and C^0 -result (Galloway-Ling-Sbierski 2018).

Splitting theorem (Beran-Ohanyan-Rott-Solis 2023)

Let (X, \ll, \leq, τ) be a globally hyperbolic LLS with global timelike $K \geq 0$. If X contains a complete timelike line (+ some technical conditions) then it splits into a product $\mathbb{R} \times S$ with S an Alexandrov space with $K \geq 0$.

Generalises smooth Lorentzian as well as synthetic Riemannian results.

More on Lorentzian (pre-)length spaces

- *causal ladder* (Kunzinger-Sämman 2018, Aké Hau-Cabrera-Solis 2020)
- *Generalized cones*, i.e., Lorentzian warped products of length spaces with 1-dim base and singularity theorems
(Alexander-Graf-Kunzinger-Sämman 2021)
- *null distance* & Lorentzian length spaces (Kunzinger-S. 2022)
- *Gluing* of Lorentzian length spaces (Beran-Rott 2022)
- Hyperbolic *angles* (Barrera-de Oca-Solis 2022, Beran-Sämman 2022)
- *time functions* on Lorentzian (pre-)length spaces
(Burtscher-García-Heveling 2021)
- Lorentzian Hausdorff *dimension, measure* (McCann-Sämman 2021)
- *Causal boundaries* (Ake Hau-Burgos-Solis 2023,
Burgos-Flores-Herrera 2023)
- *Machine learning* in spacetimes (Law-Lucas 2023)

Ricci bounds via optimal transport: the basic idea

- *Optimal Transport*: Monge, Kantorovich, move matter in the cheapest / optimal way from X to Y

- *Minimize*

$$\int_{X \times Y} c(x, y) d\pi(x, y)$$

over couplings $\pi \in \mathcal{P}(X \times Y)$ w. given marginals

$$(\text{pr}_X)_\# \pi = \mu_1, (\text{pr}_Y)_\# \pi = \mu_2$$

- What is *optimal* depends on *distances* and *geometry* !

- Turn this on its head:
define *curvature* by requiring that OT behaves as in model spaces

- Riemannian case: cost $c = d$
- Lorentzian case: cost $c = \tau$

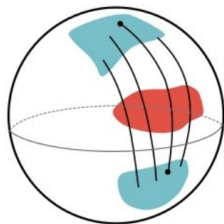


Figure: Transporting *clouds* of points on the sphere

Ricci Bounds via Optimal Transport: Riemannian case

Thm. (Ric. bds. & displacement convexity, Lott-Villani, Sturm 2006-09)

(M, g) complete Riemannian manifold

$$\text{Ric}_g \geq 0 \iff (M, d_g, \text{vol}_g) \text{ is an RCD}(0, \infty)\text{-space}$$

Definitions. On a metric measure space (X, d, \mathfrak{m}) we define

- *Wasserstein distance:* $W_2(\mu_0, \mu_1) = \left(\inf_{\pi \in \Pi} \int_{X \times X} d(x, y)^2 d\pi(x, y) \right)^{\frac{1}{2}}$
- *Wasserstein geodesic:* continuous curve $(\mu_t)_{0 \leq t \leq 1}$ in $P_2(X)$ with

$$W_2(\mu_s, \mu_t) = |t - s| \cdot W_2(\mu_1, \mu_2)$$

- *Entropy functional:* $\text{Ent}(\mu | \mathfrak{m}) = - \int \rho \log(\rho) d\mathfrak{m}$ for $\mu = \rho \mathfrak{m}$
- *RCD(0,)-space:* $\text{Ent}(\mu | \mathfrak{m})$ convex along Wasserstein geodesics

Again turn this into definition of synthetic curvature bounds.

\rightsquigarrow *theory of CD-spaces:* stability under measured GH-convergence

Ricci Bounds via Optimal Transport: Lorentzian case

Thm. (Ric. bds. & displacement conv., McCann, Mondino-Suhr 2020)

(M, g) globally hyperbolic spacetime

$\text{Ric}(X, X) \geq 0$ for X timelike $\iff (M, d_g, \text{vol}_g)$ is $\text{TCD}(K, N)$ -space

Definitions. Measured Lorentzian pre-length space $(X, d, \mathbf{m} \ll, \leq, \tau)$

- OT & causality (Eckstein-Miller 2017) $\mu_1, \mu_2 \in \Pi_{\ll}$

- *p -Lorentz Wasserstein distance:* $(0 < p < 1)$

$$l_p(\mu_1, \mu_2) = \left(\sup_{\pi \in \Pi_{\ll}} \int_{X \times X} \tau(x, y)^p d\pi(x, y) \right)^{1/p}$$

- *Entropy functional:* $\text{Ent}(\mu | \mathbf{m}) = - \int \rho \log(\rho) d\mathbf{m}$ for $\mu = \rho \mathbf{m}$

- $\text{TCD}(K, N)$: along l_p -geos μ_t we have for $e(t) := \text{Ent}(\mu_t | \mathbf{m})$

$$e''(t) - \frac{1}{N} e'(t)^2 \geq K \int_{X \times X} \tau(x, y)^2 \pi(dx dx)$$

Selected results

Hawking's singularity theorem in TMCP (Cavaletti-Mondino 2024)

Let $(X, d, \mathfrak{m} \ll, \leq, \tau)$ be a globally hyperbolic measured LLS such that

- 1 TCD(0, N) (replaces (SEC)), with
- 2 a Borel achronal FTC set V w. synthetic mean curvature $\leq H_0 < 0$.

Then $\tau_V \leq D_{H_0, 0, N}$ on $I^+(V)$.

Complements low regularity spacetime singularity theorems

(Graf 2020, Kunzinger-Ohanyan-Schinnerl-S. 2022, see S. 2023)

- Synthetic vacuum Einstein equations (Mondino-Suhr 2023)
- Differential calculus for time functions on LLS:
(Beran-Braun-Calisti-Gigli-McCann-Ohanyan-Rott-Sämman 2024)
- Lorentzian splitting (new proof for class. result, synthetic in progress)
(Braun-Gigli-McCann-Ohanyan-Sämman 2024)

Outlook

(Measured) Measured Lorentzian Length Spaces $(X, d, \mathfrak{m} \ll, \leq, \tau)$

- provide a general mathematical setting for
 - ▶ *sectional* curvature and
 - ▶ *Ricci* curvature (bounds)
- that contains
 - ▶ *low regularity spacetimes* but also
 - ▶ *discrete spaces*

Gives framework for

- approaches to non-smooth spacetime geometry
 $g \in C^{0,1}, g \in C^0 +$ causally plain
- fundamentally discrete approaches to QG
Causal set theory, Causal Fermion systems

Outlook: Causal set theory

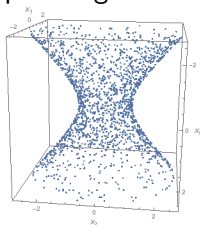
- ingredients: (X, \leq) , partial order; called *causal set* that is *locally finite*: $J(x, y) = \{z : x \leq z \leq y\}$ finite
- CS hypothesis: QT of causal sets X ; (M, g) approximation of X

$$\mathcal{C}(M, \rho_C) \ni X \longleftrightarrow (M, g)$$

Hauptvermutung of CST

X can be embedded at density ρ_C into two distinct spacetimes iff they are “close”.

sprinkling



- terminology:
 - ▶ chain: $C := (x_i)_{i=1}^n : x_i < x_{i+1}$
 - ▶ length: $L(C) = n$
 - ▶ $\tau(x, y) := \sup\{L(C) : C \text{ chain from } x \text{ to } y\}$

(X, \ll, \leq, τ) is a Lorentzian pre-length space

Hauptvermutung translates into statement on convergence of LLS.

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